

Electroweak symmetry breaking induced by dark matter

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In collaboration with Michel Tytgat: [arXiv:0707.0633](https://arxiv.org/abs/0707.0633) (hep-ph)

Motivation

- What kind of new physics can we expect at 100 GeV - 1 TeV ?
- Origin of electroweak symmetry breaking,
 - unitarity bound, triviality bound, electroweak precision constraints,...
 - the Dark Matter particle
 - from galaxy rotation curves, missing mass in clusters, lensing, CMB and LSS, bullet clusters... $\Rightarrow \Omega_{DM} \simeq 0.23$
 - If $\sigma_{annihil.} \sim g^4/m_{DM}^2$, $\Omega_{DM} \sim 0.23$
requires $m_{DM} \sim 100$ GeV (WIMP mechanism)
 - Hierarchy problem(s): suggest new low scale physics, SUSY,...
- ⇒ ⇒ Could these issues be related???

Relation between EWSB and DM?

 This talk: these 2 issues could be deeply related



A model where EWSB is due to the existence of DM, a la
Coleman-Weinberg

The model

 very minimal assumptions:

- 2 Higgs doublets: H_1, H_2
- a Z_2 symmetry: all particles even, except H_2 which is odd

 the lightest H_2 component is stable, i.e. a DM candidate



“Inert Higgs doublet model”: H_2 has no vev, no mixing with H_1 and no couplings to SM fermions

Deshpande, Ma 78'; Ma 06'; Barbieri, Hall, Rychkov 06',...

Dark Matter

- $H_2 = \begin{pmatrix} H^+ \\ \frac{H_0 + iA_0}{\sqrt{2}} \end{pmatrix} \longrightarrow$ 4 components: H^\pm, H_0, A_0

 if H_0 or A_0 is the lightest state: good DM candidate

- Most general scalar potential:

$$V = \mu_1^2 |H_1|^2 + \mu_2^2 |H_2|^2 + \lambda_1 |H_1|^4 + \lambda_2 |H_2|^4 \\ + \lambda_3 |H_1|^2 |H_2|^2 + \lambda_4 |H_1^\dagger H_2|^2 + \frac{\lambda_5}{2} \left[(H_1^\dagger H_2)^2 + h.c. \right]$$

Dark Matter: tree level mass spectrum

$$m_{H^+}^2 = \mu_2^2 + \lambda_3 \frac{v^2}{2}$$

$$m_{H_0}^2 = \mu_2^2 + \lambda_L \frac{v^2}{2} \quad (\lambda_L = \lambda_3 + \lambda_4 + \lambda_5)$$

$$m_{A_0}^2 = \mu_2^2 + \lambda_S \frac{v^2}{2} \quad (\lambda_S = \lambda_3 + \lambda_4 - \lambda_5)$$

 To have H_0 or A_0 the lightest: $\lambda_L < \lambda_3$ or $\lambda_S < \lambda_3$

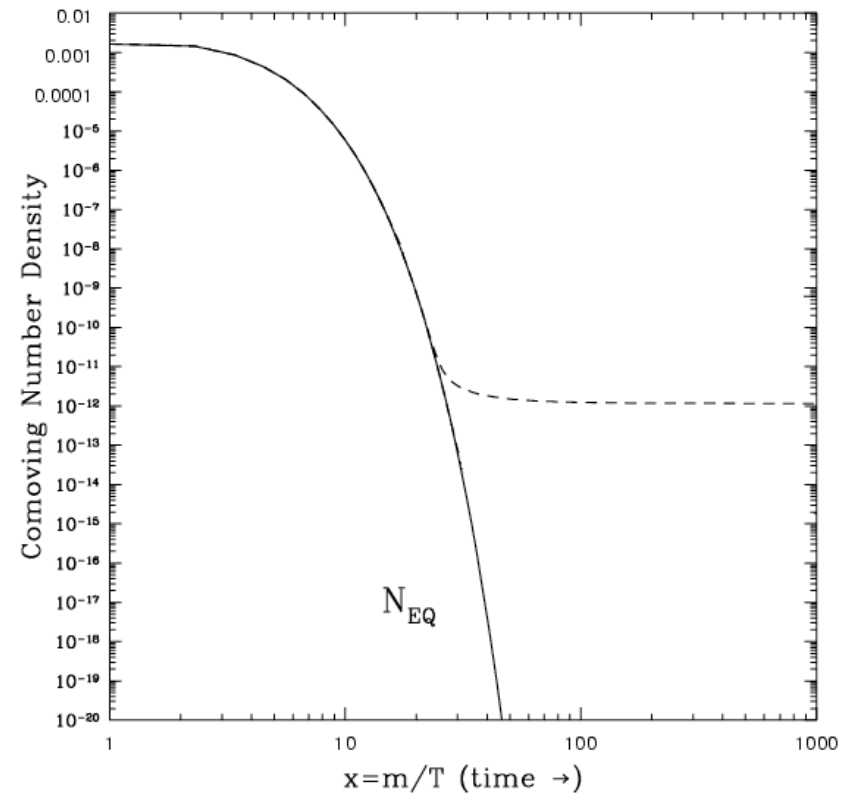
Relic DM density: annihilation freeze out mechanism

- Down to $T \sim m_{DM}$, DM is in thermal equilibrium: $n_{DM} \simeq n_{DM}^{Eq}$
- For $T < m_{DM}$: $n_{DM} \propto e^{-m_{DM}/T}$ ← Boltzmann suppression

freeze out of the annihilation at $T = T_f$

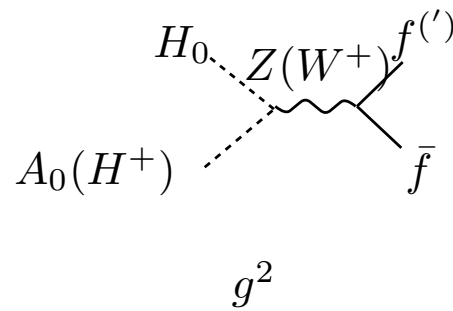
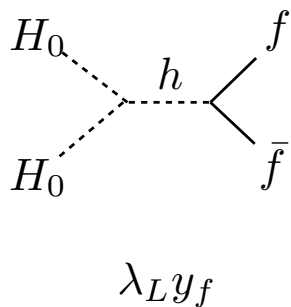
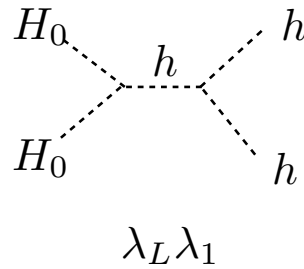
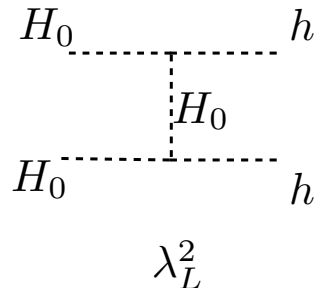
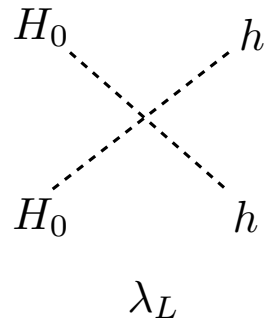
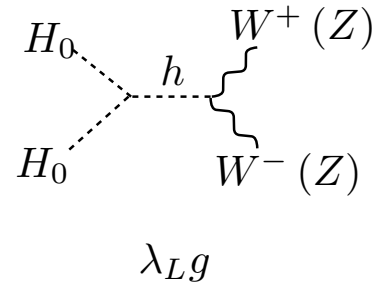
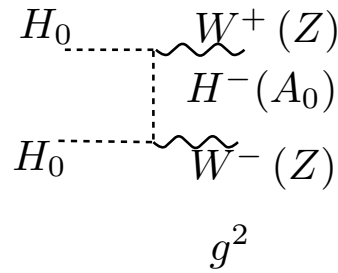
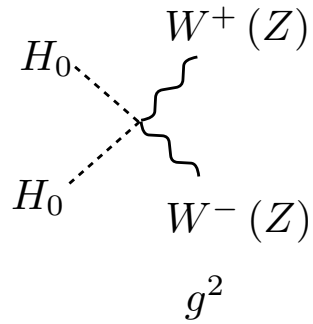
$$n_{DM}(T < T_f) = n_{DM}^{Eq}(T = T_f)$$

⇒ If $\sigma_{annihil.} \sim g^4/m_{DM}^2$, $\Omega_{DM} \sim 0.23$
requires $m_{DM} \sim 100$ GeV (WIMP mechanism)



Relic DM density: annihilation processes

If H_0 is the DM:



Two possible DM mass regimes

- Low DM mass regime: $m_{DM} < m_W$ ($DM = H_0$ or A_0)

→ To avoid too fast $DM DM \rightarrow W W, Z Z$ annihilation

$$40 \text{ GeV} < m_{DM} < 75 \text{ GeV}$$



Detailed analysis: Lopez-Honorez, Nezri, Oliver, Tytgat '06;
see also Barbieri, Hall, Rychkov '06;
Bergstrom et al '07




In agreement with direct and indirect
gamma detection constraints

Two possible DM mass regimes

- High DM mass regime: $m_{DM} > 400 \text{ GeV}$

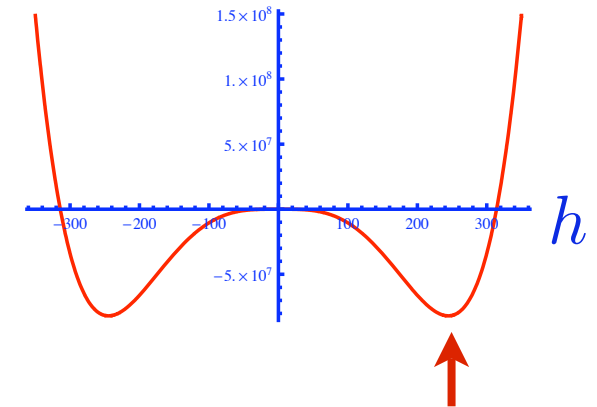
 Not relevant for EWSB a la Coleman-Weinberg


DM couplings too large

EWSB: Coleman-Weinberg mechanism in the SM

- Standard Model: $V = \mu^2 H^2 + \lambda H^4$

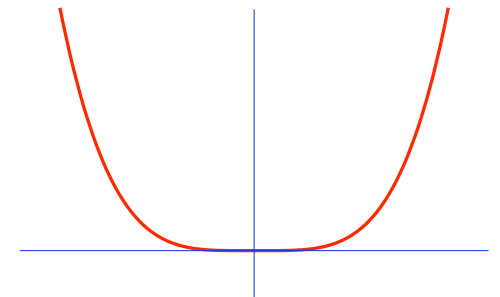
⇒ only one scale: μ^2



$$\langle h \rangle = v = 246 \text{ GeV}$$

- Coleman-Weinberg: could we start with no scale at all or at least no large scale at all? ← $\mu^2 \simeq 0$

$$V \simeq \lambda H^4$$

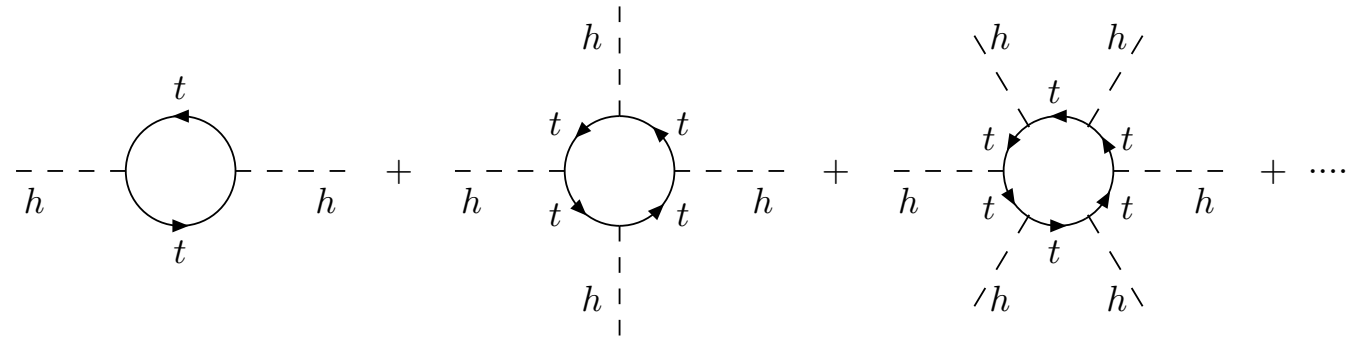


no EWSB at tree level

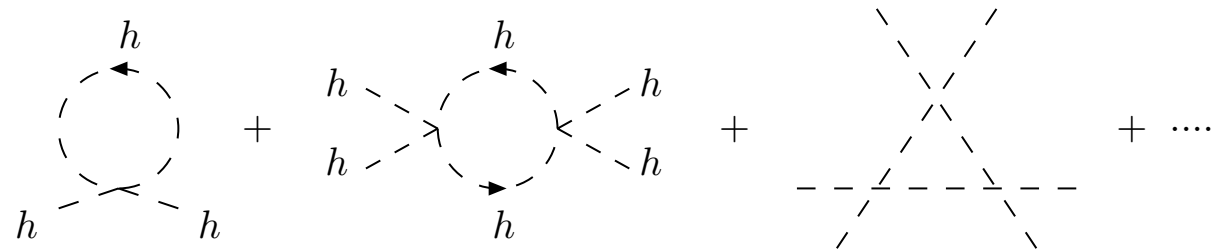
Coleman-Weinberg in the SM: one loop diagrams

S. Coleman, E. Weinberg '73

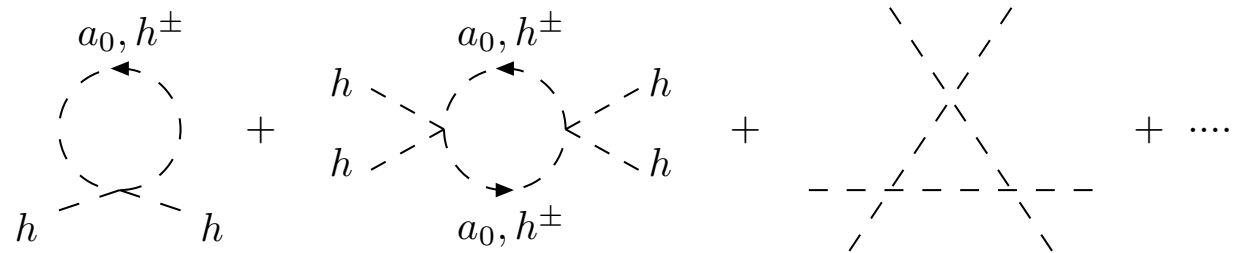
Top loops:



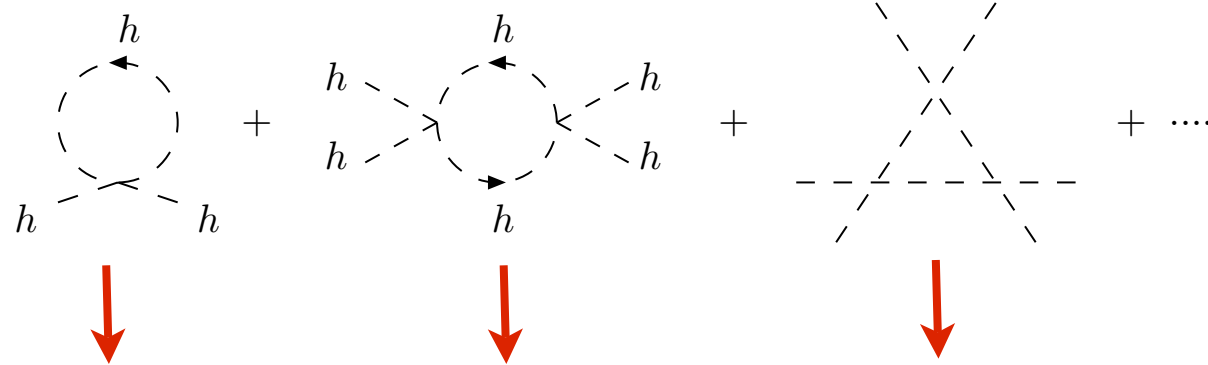
Higgs loops:



Goldstone loops:



Coleman-Weinberg in the SM: summation of diagrams



$$V_{eff}(h) \ni i \int \frac{d^4k}{(2\pi)^4} \frac{1}{2} \left(3 \frac{\lambda h^2}{k^2 + i\epsilon} + \frac{3^2}{2} \frac{\lambda^2 h^4}{(k^2 + i\epsilon)^2} + \frac{3^3}{3} \frac{\lambda^3 h^6}{(k^2 + i\epsilon)^3} + \dots \right)$$

$$= \frac{1}{2} \int \frac{d^4k}{(2\pi)^4} \log \left(1 + \frac{\lambda h^2}{2k^2} \right)$$

$$= \frac{9}{64\pi^2} \lambda^2 h^4 \left(\log \frac{3\lambda h^2}{\mu^2} - \frac{3}{2} \right) \longleftarrow \overline{MS} \text{ scheme}$$

renormalization scale

Full effective potential in the SM

$$\begin{aligned} V_{\text{eff}}(h) &= \lambda_1 \frac{h^4}{4} && \leftarrow \text{Tree level} \\ &+ \frac{1}{64\pi^2} 9\lambda_1^2 h^4 \left(\ln \frac{3\lambda_1 h^2}{\mu^2} - \frac{3}{2} \right) && \leftarrow \text{Higgs loops} \\ &+ \frac{3}{64\pi^2} \lambda_1^2 h^4 \left(\ln \frac{\lambda_1 h^2}{\mu^2} - \frac{3}{2} \right) && \leftarrow \text{Goldstone loops} \\ &- \frac{12}{64\pi^2} \frac{g_t^4}{4} h^4 \left(\ln \frac{g_t^2 h^2}{2\mu^2} - \frac{3}{2} \right) && \leftarrow \text{top loops} \end{aligned}$$

\Rightarrow Imposing $\left. \frac{dV_{\text{eff}}}{dh} \right|_{h=246 \text{ GeV}} = 0$ determines $\lambda = \text{fct}(g_t)$

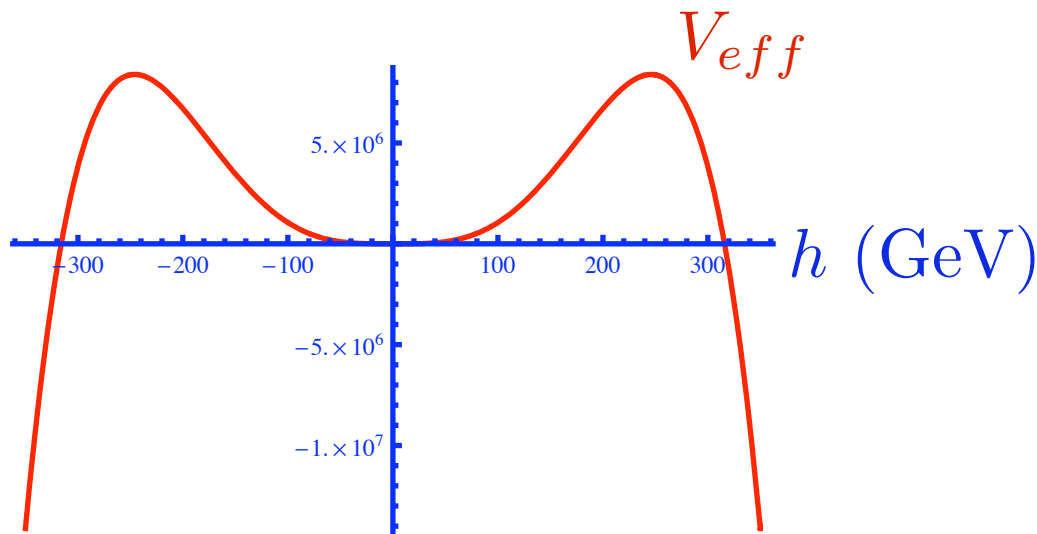
\Rightarrow No free parameter: potential known

Coleman-Weinberg in the SM: final result

↪ dominated by the top loop contribution



has the wrong sign

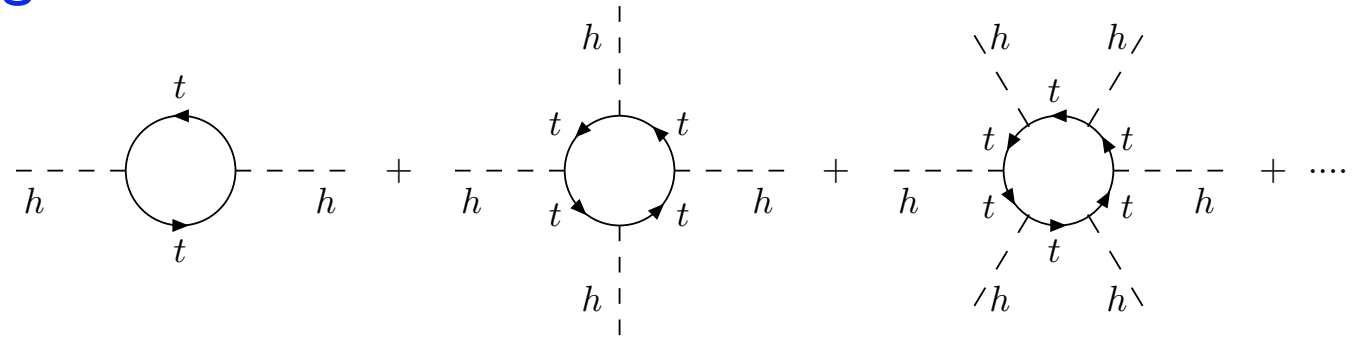


⇒ Coleman-Weinberg doesn't work in the Standard Model

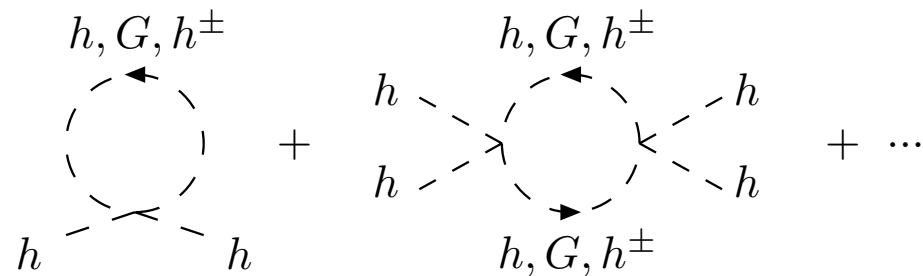
Coleman-Weinberg with an additional inert Higgs doublet: one loop diagrams

- Tree level: no scale: $\mu_1 \simeq \mu_2 \simeq 0$ (no EWSB)
- One loop diagrams:

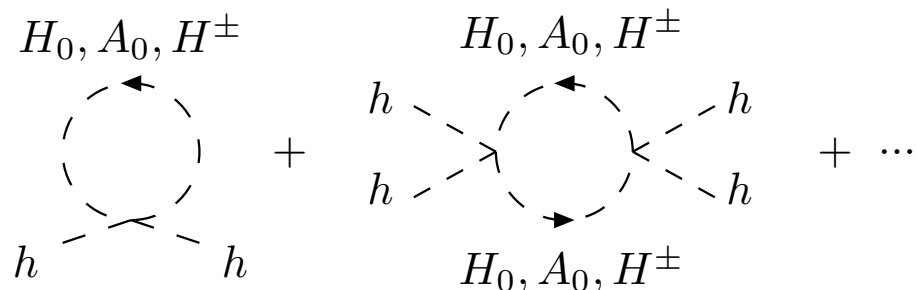
Top loops:



Higgs loops:



Inert Higgs loops:



Effective potential with a inert Higgs

- Total one loop effective potential:

$$\begin{aligned}
 V_{\text{eff}}(h) &= \lambda_1 \frac{h^4}{4} && \leftarrow \text{Tree level} \\
 &+ \frac{1}{64\pi^2} 9\lambda_1^2 h^4 \left(\ln \frac{3\lambda_1 h^2}{\mu^2} - \frac{3}{2} \right) && \leftarrow \text{Higgs loops} \\
 &+ \frac{3}{64\pi^2} \lambda_1^2 h^4 \left(\ln \frac{\lambda_1 h^2}{\mu^2} - \frac{3}{2} \right) && \leftarrow \text{Goldstone loops} \\
 &- \frac{12}{64\pi^2} \frac{g_t^4}{4} h^4 \left(\ln \frac{g_t^2 h^2}{2\mu^2} - \frac{3}{2} \right) && \leftarrow \text{top loops} \\
 &+ \frac{2}{64\pi^2} \frac{\lambda_3^2}{4} h^4 \left(\ln \frac{\lambda_3 h^2}{2\mu^2} - \frac{3}{2} \right) && \leftarrow H^\pm \text{ loops} \\
 &+ \frac{1}{64\pi^2} \frac{\lambda_L^2}{4} h^4 \left(\ln \frac{\lambda_L h^2}{2\mu^2} - \frac{3}{2} \right) && \leftarrow H_0 \text{ loops} \\
 &+ \frac{1}{64\pi^2} \frac{\lambda_S^2}{4} h^4 \left(\ln \frac{\lambda_S h^2}{2\mu^2} - \frac{3}{2} \right) && \leftarrow A_0 \text{ loops}
 \end{aligned}$$

The terms are grouped into two categories:

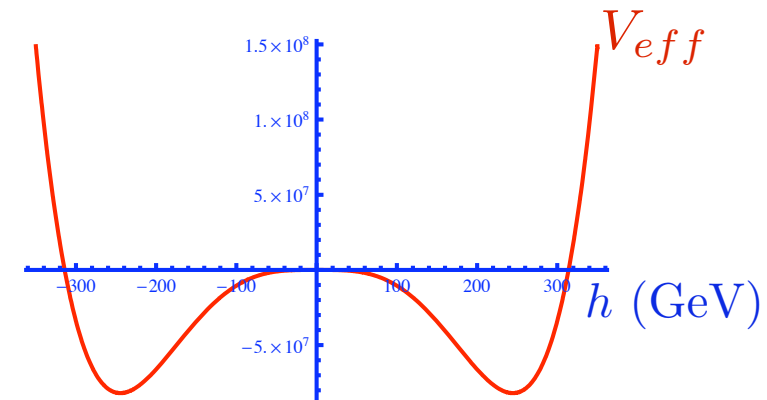
- SM contribution:** Tree level, Higgs loops, Goldstone loops, and top loops.
- new inert Higgs contribut.:** H^\pm loops, H_0 loops, and A_0 loops.

EWSB: Inert Higgs contribution

→ a scalar contribution: has the right sign

Coleman, Weinberg '73; Gildener, Weinberg '76; Meissner, Nicolai '07; Espinosa, Quiros '07; Foot et al. '07

⇒ all we need is the inert Higgs contribution larger than the top one (λ_3 and/or λ_L and/or λ_S large)



- to have a maximum we need one quartic coupling to be larger than $\sim 3g_t^2$
- to have $m_H > 114.4$ GeV we need one quartic coupling to be larger than $\sim 6g_t^2$

⇒ requires fairly large quartic coupling(s) (still perturbative)

EWSB vs DM constraints

- EWSB: at least one among $\lambda_{3,L,S}$ must be large (i.e. at least one H^+ , H_0 , A_0 must be heavy, above ~ 350 GeV)
- DM: H_0 or A_0 must be below the W (i.e. λ_L or λ_S smaller)

⇒ Large mass splittings are required inside the inert Higgs

↪ Electroweak precision measurement??



T parameter constraint

T parameter constraint

$$T = \frac{\Pi_{WW}(0)}{M_W^2 \alpha} - \frac{\Pi_{ZZ}(0)}{M_Z^2 \alpha}$$

→ Sensitive to mass splittings between members of a SU(2) multiplet

However: vanishes if $m_{H^+} - m_{A_0}$ or $m_{H^+} - m_{H_0}$ vanishes



custodial symmetry



twisted custodial symmetry

⇒ to get $|\Delta T| < 0.2$ we need: $|m_{H^+} - m_{A_0}| < \sim 40 \text{ GeV}$
or $|m_{H^+} - m_{H_0}| < \sim 40 \text{ GeV}$

→ for example: $m_{H_0} = 70 \text{ GeV}$, $m_{H^+} = 110 \text{ GeV}$, $m_{A_0} = 350 \text{ GeV}$

Custodial symmetry

→ SO(3) symmetry: nice way to justify:

- necessary mass spectrum
- T parameter constraint
- no CP violation (i.e. no phase in λ_5 , otherwise bad for DM)

→ Either with normal custodial symmetry ($m_{A_0} \sim m_{H^+}$) or with twisted custodial symmetry ($m_{H_0} \sim m_{H^+}$)

↓
Gerard, Herquet '07

Examples of sets of parameters

$$\Omega_{DM}h^2 = 0.11$$

λ_1	λ_2	λ_3	λ_4	λ_5	M_h	M_{H_0}	M_{A_0}	M_{H^\pm}	h_{BR}	W_{BR}
-0.11	0	5.4	-2.8	-2.8	120	12	405	405	100%	0%
-0.11	-2	5.4	-2.7	-2.7	120	43	395	395	100%	0%
-0.11	-3	5.4	-2.6	-2.6	120	72	390	390	94%	6 %
-0.30	0	7.6	-4.1	-4.1	180	12	495	495	100%	0 %
-0.30	-2.5	7.6	-3.8	-3.8	180	64	470	470	100%	0 %
-0.18	-3	-0.003	4.6	-4.7	120	39	500	55	100%	0 %
-0.29	-5	-0.07	5.5	-5.53	150	54	535	63	0%	100 %

⇒ possible range of m_{DM} extended: $10 \text{ GeV} < m_{DM} < 75 \text{ GeV}$

More general situation: large μ^2

→ In this model the DM Higgs doublet can have dramatic effect on EWSB parameters even for large μ^2 as soon as some of the quartic couplings are sizeable

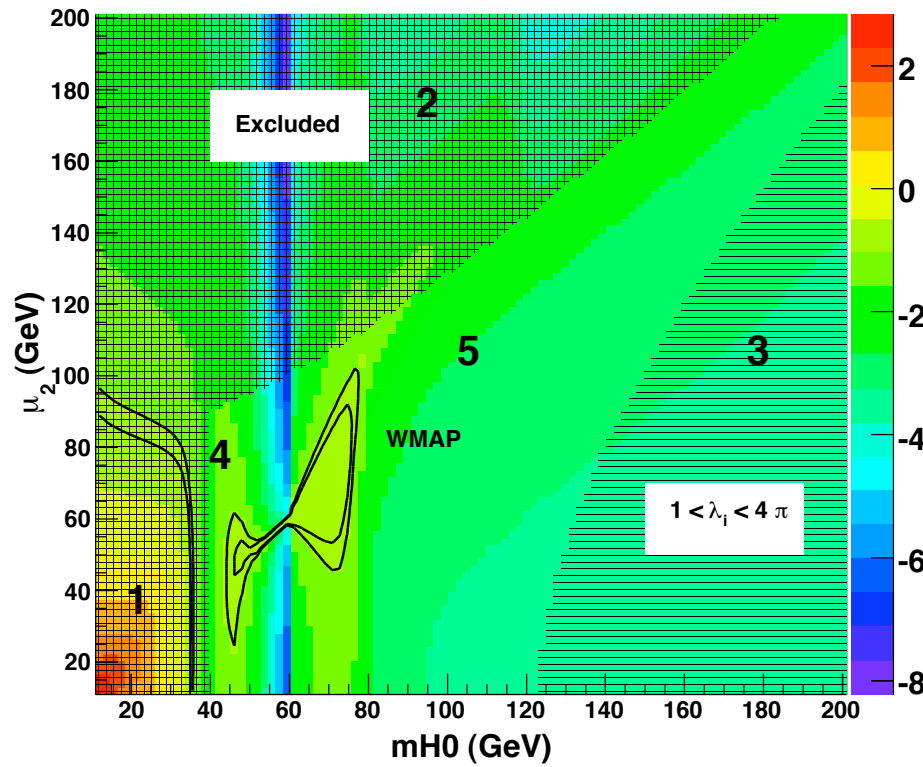
→ Might be generic for other models

Summary

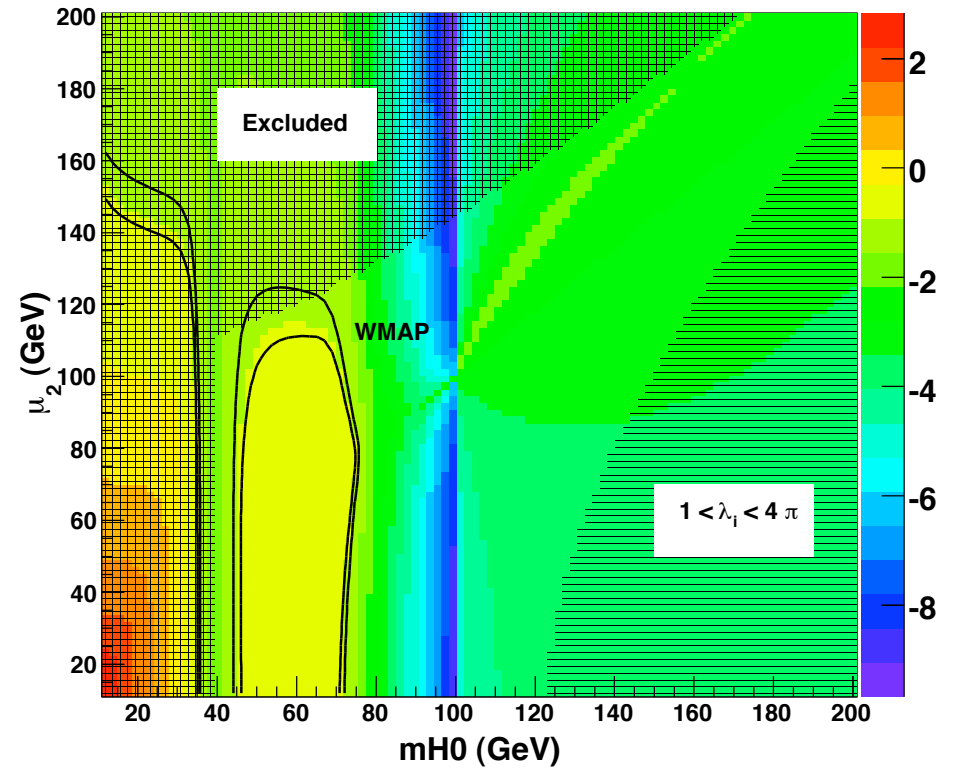
- EWSB and WIMP DM: similar scales \Rightarrow possible link???
- DM could be crucial for EWSB: explicit minimal example where a single additional Higgs doublet with Z_2 parity provides a good DM candidate and allows EWSB a la Coleman-Weinberg
- In this framework the DM mass is proportional to the EW scale \Rightarrow provides a possible hint for why DM would be at the EW scale as required by the WIMP mechanism

Relic density

$\log_{10} [\Omega h^2] : m_h=120 \text{ GeV} ; I_2=10^{-1} ; \Delta MA_0= 10 \text{ GeV} ; \Delta MH_c= 50 \text{ GeV}$



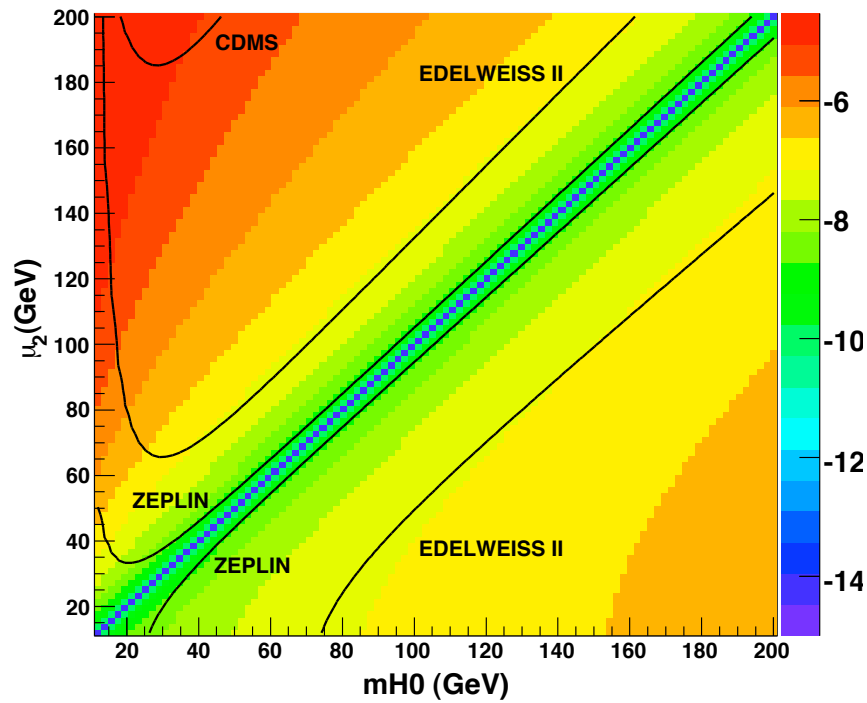
$\log_{10} [\Omega h^2] : m_h=200 \text{ GeV} ; I_2=10^{-1} ; \Delta MA_0= 10 \text{ GeV} ; \Delta MH_c= 50 \text{ GeV}$



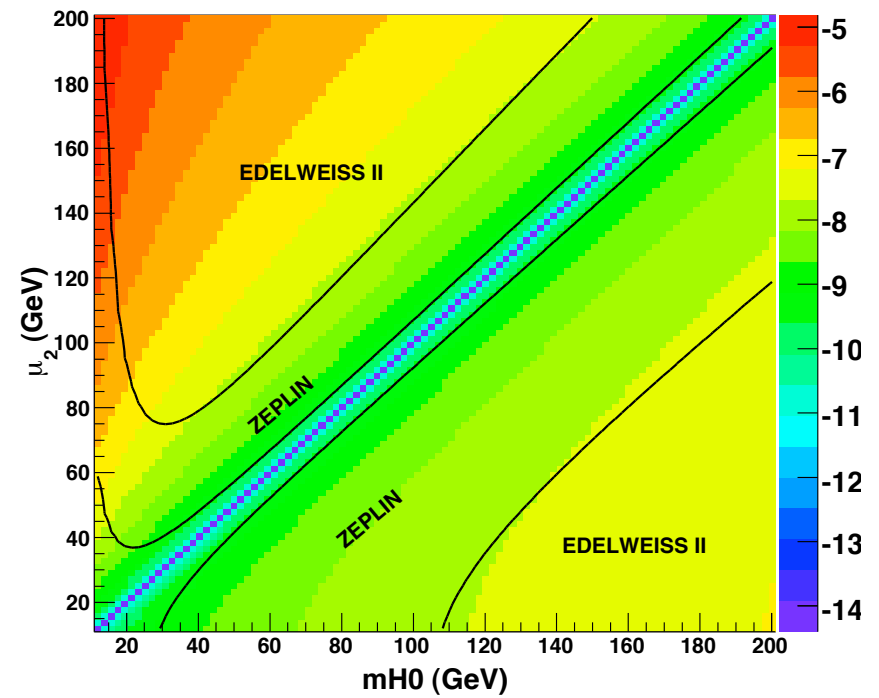
Lopez-Honorez, Nezri, Oliver, Tytgat '06

Direct detection

$\log_{10} [\sigma_{\text{DM-p}}]$: $m_h=120$ GeV ; $I_2=10^{-1}$; $\Delta \text{MA0}=10$ GeV ; $\Delta \text{MHc}=50$ GeV



$\log_{10} [\sigma_{\text{DM-p}}]$: $m_h=200$ GeV ; $I_2=10^{-1}$; $\Delta \text{MA0}=10$ GeV ; $\Delta \text{MHc}=50$ GeV



Indirect detection

$\log_{10} [\text{flux}_{1\gamma} (\text{cm}^{-2}\text{s}^{-1})]$: $m_h=120 \text{ GeV}$; $I_2=10^{-1}$; $\Delta M_{A0}=10 \text{ GeV}$; $\Delta M_{Hc}=50 \text{ GeV}$

