GRAND UNIFICATION WITH & WITHOUT SUPERSYMMETRY

Alejandra Melfo Centro de Física Fundamental Mérida, Venezuela

IN COLLABORATION WITH.

Goran Senjanovic (ICTP) Borut Bajc (Ljubljana) Francesco Vissani (Gran Sasso) Alba Ramírez (Mérida)

BW2007, KLADOVO, SEP. 2-9, 2007

OUTLINE

2

- GUTs and neutrino mass
- Intermediate scales
- The see-saw and the Yukawa sector
- Non supersymmetric models
- The minimal SUSY model
- Departing from the minimal



GUTS AND NEUTRINO MASS

SO(10): all fermions in <u>16</u> representation

SU(5) fermions: in <u>5</u> and <u>10</u> representations

 $\Rightarrow \nu_R$ is a singlet

- adding a singlet to the theory gives a lot of new parameters
- SU(5) breaks directly to SU(3)xSU(2)xU(1)

- no intermediate scales

... and $m_{
u}$ calls for intermediate scales

THE (B-L) BREAKING SCALE

Best idea for small m_{ν} : the see-saw mechanism

give ν_R a mass by breaking B-L at a large scale M_R

$$\langle \Delta \rangle \nu_R^T i \sigma_2 \nu_R \qquad \langle \Delta \rangle = M_R$$

$$m_{\nu} = \frac{M_W^2}{M_R} \qquad m_{\nu}$$

$$m_{\nu} \sim 0.01 eV$$

 $M_R \sim 10^{13} GeV$

An intermediate scale would be convenient (not indispensable)



SUSY: ONE-STEP UNIFICATION

$$\frac{1}{\alpha_i(M_W)} = \frac{1}{\alpha_U} - \frac{b_i}{2\pi} \ln(M_G/M_W)$$

5

 $M_G \sim 10^{16} GeV$

NON-SUSY: INTERMEDIATE SCALE

$$\frac{1}{\alpha_i(M_W)} = \frac{1}{\alpha_U} - \frac{b_i}{2\pi} \ln(M_R/M_W)$$
$$-\frac{b'_i}{2\pi} \ln(M_G/M_R)$$

 M_R determined by the particle content

SO(10) SYMMETRY BREAKING

6

7.6 .1.1	SO(10)	
Many possible intermediate scales	$M_X\Downarrow\langle p angle$	GUT scale
	$SU(4)_C \times SU(2)_L \times SU(2)$	R
	$M_{PS} \Downarrow \langle a \rangle$	
SU	$T(3)_C \times SU(2)_L \times SU(2)_R \times$	$U(1)_{B-L}$
	$M_R\Downarrow \langle \sigma angle$	see-saw scale
	$SU(3)_C \times SU(2)_L \times U(1)_2$	Y
A. MELFO		

TWO TYPES OF SEE-SAW

7

TYPE I (renormalizable version)

• An $SU(2)_R$ triplet with (B - L) = 2 gets a vev at a large scale M_R

 $\langle \Delta^c \rangle \Rightarrow \nu^c \text{ mass} \sim M_R$

gives a mass to the right-handed neutrino

• At EW scale, neutrino gets a Dirac mass m_D

$$\begin{pmatrix} 0 & m_D \\ m_D & M_R \end{pmatrix} \longrightarrow m_\nu \sim \frac{m_D^2}{M_R} \sim \frac{M_W^2}{M_R}$$

TYPE II

In Left-Right theories, terms like:

 $\Delta H^2 \Delta^c + m_{\Delta} \Delta^2$

H: bidoublet Δ : Left-handed triplet Δ^c : Right-handed triplet

Provide a small vev for the Left-handed triplet after EW breaking

 $\langle \Delta \rangle \sim \frac{\langle H \rangle^2 \langle \Delta^c \rangle}{m^2} \sim \frac{M_W^2}{M_D}$ Mass for ν from $L^T \tau_2 \langle \Delta \rangle L$

vev of Δ^c induces a small vev for Δ after EW breaking

In SUSY SO(10), triplets are in 126: mixing with 54 or 210 can give such terms in the potential.

> TWO TYPES OF SEE-SAW ARE OF SAME MAGNITUDE: BUT VERY DIFFERENT PARAMETERS INVOLVED



SU(4)c X SU(2)L X SU(2)R DECOMPOSITION

10

 $H_{10} = (6,1,1) + (1,2,2)$ $D_{120} = (\overline{10},1,1) + (10,1,1) + (6,3,1) + (6,1,3) + (1,2,2) + (15,2,2)$ $\overline{\Sigma}_{\overline{126}} = (10,1,3) + (\overline{10},3,1) + (6,1,1) + (15,2,2)$ $\Delta_R \qquad \Delta_L$

- 126 can give type I and type II see-saw
- (15,2,2) in 126 can contain the SM Higgs

A. MELFO

is 126 enough for all fermion masses ? no..

One doublet is not enough:

Lazarides, Shafi Wetterich 1981

Clark, Kuo Nakagawa 1982

 $M_U = y_{10} \langle 1, 2, 2 \rangle_{10}^u + y_{126} \langle 15, 2, 2 \rangle_{126}^u$ $M_D = y_{10} \langle 1, 2, 2 \rangle_{10}^d + y_{126} \langle 15, 2, 2 \rangle_{126}^d$ $M_E = y_{10} \langle 1, 2, 2 \rangle_{10}^d - 3 y_{126} \langle 15, 2, 2 \rangle_{126}^d$

- only 10: $m_d = m_l$ • only 126: $3 m_d = m_l$ } at the GUT scale, for all generations
- 126 required for neutrino mass but what else?
 - is there a difference between choosing 10 or 120 ?

Notice: same question for SUSY or non-SUSY models

NON-SUSY: 126 + 10

(2nd and 3rd generations only) Ba

Bajc, A.M, Senjanovic, Vissani 2005

 $M_{U} = y_{10} \langle 1, 2, 2 \rangle_{10}^{u} + y_{126} \langle 15, 2, 2 \rangle_{126}^{u}$ $M_{D} = y_{10} \langle 1, 2, 2 \rangle_{10}^{d} + y_{126} \langle 15, 2, 2 \rangle_{126}^{d}$ $M_{E} = y_{10} \langle 1, 2, 2 \rangle_{10}^{d} - 3 y_{126} \langle 15, 2, 2 \rangle_{126}^{d}$ $M_{\nu_{D}} = y_{10} \langle 1, 2, 2 \rangle_{10}^{u} - 3 y_{126} \langle 15, 2, 2 \rangle_{126}^{u}$

 $M_{\nu_L} = y_{126} \langle \overline{10}, 3, 1 \rangle_{126}^d$ $M_{\nu_R} = y_{126} \langle 10, 1, 3 \rangle_{126}^d$

12

see-saw, type I and II: $M_N = -M_{\nu_D} M_{\nu_R}^{-1} M_{\nu_D} + M_{\nu_L}$

approx.
$$\theta_q = V_{cb} = 0$$

 $\frac{\langle 2, 2, 1 \rangle_{10}^u}{\langle 2, 2, 1 \rangle_{10}^d} = \frac{m_c(m_\tau - m_b) - m_t(m_\mu - m_s)}{m_s m_\tau - m_\mu m_b} \approx \frac{m_t}{m_b}$

• real 10: $m_t = m_b$

SUSY OR NOT: 126 + 10

Bajc, Senjanovic, Vissani 2002

13

 $M_D = y_{10} \langle 1, 2, 2 \rangle_{10}^d + y_{126} \langle 15, 2, 2 \rangle_{126}^d$ $M_E = y_{10} \langle 1, 2, 2 \rangle_{10}^d - 3 y_{126} \langle 15, 2, 2 \rangle_{126}^d$

Type II see-saw: $M_N = M_{\nu_L} = y_{126} \langle 10, 1, 3 \rangle_{126}^d$

$$\theta_D = 0 \text{ (small mixing in } M_D) \qquad \qquad M_N \propto \begin{pmatrix} 0 & 0 \\ 0 & m_b - m_\tau \end{pmatrix}$$

unless $m_b = m_{\tau}$, neutrino mixing vanishes

large
$$\theta_{atm} \leftrightarrow b - \tau$$
 unification

Full 3-gen. analysis:

- connection still true
- - θ_{13} close to exp. limit

Matsuda, Koide, Fukuyama, Nishiura 2002

Goh, Mohapatra, Ng, 2003

NON-SUSY: 126 + 120

14

(2nd and 3rd generations only)

 $M_{U} = y_{120}(\langle 1, 2, 2 \rangle_{120}^{u} + \langle 15, 2, 2 \rangle_{120}^{u}) + y_{126} \langle 15, 2, 2 \rangle_{126}^{u}$ $M_{D} = y_{120}(\langle 1, 2, 2 \rangle_{120}^{d} + \langle 15, 2, 2 \rangle_{120}^{d}) + y_{126} \langle 15, 2, 2 \rangle_{126}^{d}$ $M_{E} = y_{120}(\langle 1, 2, 2 \rangle_{120}^{d} - 3 \langle 15, 2, 2 \rangle_{120}^{d}) - 3 y_{126} \langle 15, 2, 2 \rangle_{126}^{d}$ $M_{\nu_{D}} = y_{120}(\langle 1, 2, 2 \rangle_{120}^{u} - 3 \langle 15, 2, 2 \rangle_{120}^{u}) - 3 y_{126} \langle 15, 2, 2 \rangle_{126}^{u}$ $y_{120} \text{ antisymmetric}$

• real 120: $m_t = m_b$

complex 120: interesting connections with neutrino masses and mixings

15 SUSY OR NOT: 126 + 120 (2nd and 3rd generations only) Bajc, A.M, Senjanovic, Vissani 2005 Defining some small ratios : $\epsilon_f = m_2^f / m_3^f$ • neutrino masses: $\frac{m_3^2 - m_2^2}{m_2^2 + m_2^2} = \frac{\cos 2\theta_A}{1 - \sin^2 2\theta_A/2} + \mathcal{O}(|\epsilon|)$ PREDICTIONS: • large θ_A gives degenerate neutrinos quark masses relation at the GUT scale: $\frac{m_{\tau}}{m_{\tau}} = 3 + 3\sin 2\theta_A \operatorname{Re}[\epsilon_E - \epsilon_D] + \mathcal{O}(|\epsilon^2|)$ • but $m_{\tau} \approx 2m_b$. - include running, 3-gen. effects... • quark mixing: $|V_{cb}| \simeq \cos 2\theta_A \frac{m_s}{m_b} + O(|\epsilon^2|)$ large neutrino mixing implies small quark mixing A. MELFO

126 + 10 OR 126 + 120?

- In non-supersymmetric models, both posible in principle
 - 10 and 120 need to be complex
 - can have a PQ symmetry axion as DM
- SUSY requires 126 + 10 for $m_b = m_{\tau}$
- Type II (even Type I) see-saw can give relations between neutrino and charged fermions masses and mixings
- Detailed models can be even more predictive: symmetry breaking and unification constraints
- 10 + 120 ? radiative see-saw works for split-SUSY

Bajc, Senjanovic 2005

16

		12
	UNIFICATION:	
	NON-SUSY Deshpande	, Keith, Pal 1993
$m_{\nu} \ge$	$m_t^2/M_R \longrightarrow M_R \ge 10^{13} \text{ GeV}$	$\log(M_R/GeV)$
I :	$SO(10) \longrightarrow \{2_L 2_R 4_C\} \longrightarrow \{2_L 2_R 1_X 3_c\} \longrightarrow \{2_L 1_Y 3_c\}$	8.2-10.6
II :	$SO(10) \xrightarrow[54]{} \{2_L 2_R 4_C P\} \xrightarrow[210]{} \{2_L 2_R 1_X 3_c P\} \xrightarrow[h]{} \{2_L 1_Y 3_c\}$	8.6 - 13.6
III :	$SO(10) \xrightarrow{54} \{2_L 2_R 4_C P\} \xrightarrow{45} \{2_L 2_R 1_X 3_c\} \xrightarrow{h} \{2_L 1_Y 3_c\}$	8.0 - 13.6
IV :	$SO(10) _{14} \{2_L 2_R 1_X 3_c P\} _{210} \{2_L 2_R 1_X 3_c\} _{h} \{2_L 1_Y 3_c\}$	8.2-10.8
V :	$SO(10) \xrightarrow{210} \{2_L 2_R 4_C\} \xrightarrow{15} \{2_L 1_R 4_C\} \xrightarrow{1} \{2_L 1_Y 3_C\}$	11.0-11.2
VI :	$SO(10) \xrightarrow{54} \{2_L 2_R 4_C P\} \xrightarrow{45} \{2_L 1_R 4_C\} \xrightarrow{h} \{2_L 1_Y 3_c\}$	12.2 - 13.6
VII :	$SO(10) \xrightarrow{54} \{2_L 2_R 4_C P\} \xrightarrow{210} \{2_L 2_R 4_C\} \xrightarrow{h} \{2_L 1_Y 3_c\}$	11.3 - 13.6
VIII :	$SO(10) \longrightarrow \{2_L 2_R 1_X 3_c\} \longrightarrow \{2_L 1_R 1_X 3_c\} \longrightarrow \{2_L 1_Y 3_c\}$	2.0.7.7
IX :	$SO(10) {} \{2_L 2_R 1_X 3_c P\} {} \{2_L 1_R 1_X 3_c\} {} \{2_L 1_Y 3_c\}$	2.0.10.0
X :	$SO(10) {} \{2_L 2_R 4_C\} {} \{2_L 1_R 1_X 3_c\} {} \{2_L 1_V 3_c\}$	
XI :	$SO(10) \longrightarrow \{2_L 2_R 4_C P\} \xrightarrow{210} \{2_L 1_R 1_X 3_c\} \longrightarrow \{2_L 1_V 3_c\}$	0.040.50
XII :	$SO(10) \longrightarrow \{2_L 1_R 4_C\} \longrightarrow \{2_L 1_R 1_X 3_c\} \longrightarrow \{2_L 1_V 3_c\}$	2.0.5.3

UNIFICATION: SUSY

- One-step: no intermediate scales
 - $m_{\nu} \propto M_W^2/M_{GUT}$ can be too small
- Potentials very constrained: no survival principle
 - calculate all the masses
- See-saw + SUSY = MSSM with R-parity
 - R-parity is in the center of SO(10)
 - R-parity \equiv Matter parity = $(-1)^{3(B-L)}$
 - See-saw: break (B-L) with a (B-L)-even field in order to give v_R a mass ...get R-parity preserved and the stable LSP is a DM candidate

Aulakh, A.M, Rasin, Senjanovic 1998

18

WHAT IS THE MINIMAL RENORMALIZABLE SUSY- GUT ?

- Based on SO(10)
- With a see-saw for neutrino mass: $\overline{126}$ (+ 126)
- Yukawa sector: $10 + \overline{126}$ needed: the light Higgs must be a combination of doublets in 10 and $\overline{126}$
 - need a mixing $\langle \Phi \rangle H_{10} \overline{\Sigma}_{\overline{126}}$ can use 210
- Symmetry breaking down to LR (126, 126 break down to MSSM)

Babu, Mohapatra, 1993

19

 $\Phi_{210}, H_{10}, \overline{\Sigma}_{126}, \Sigma_{126}$

210 can do that too

MINIMAL SO(10)

Clark, Kuo, Nakagawa,1982

20

Aulakh, Bajc, A.M, Senjanovic, Vissani,2003

 $\Psi_{16}, H_{10}, \Sigma_{126}, \overline{\Sigma}_{\overline{1}26}, \Phi_{210}$

 $W_{H} = m_{\Phi} \Phi^{2} + m_{\Sigma} \Sigma \overline{\Sigma} + \lambda \Phi^{3} + \eta \Phi \Sigma \overline{\Sigma} + m_{H} H^{2} + \Phi H (\alpha \Sigma + \bar{\alpha} \overline{\Sigma})$ + $y_{10} \Psi C \Gamma \Psi H + y_{126} \Psi C \Gamma^{5} \Psi \overline{\Sigma}$

- 26 real parameters: same as MSSM
- light Higgs made up of 126, 10 and 210 doublets
 - rich enough Yukawa structure
- Type I and II see-saw
 - possibility of connecting large θ_A with $b \tau$ unification
- symmetry can be broken down to MSSM (+R-parity)
 - stable LSP



• Find the composition of the light Higgs doublets

Aulakh, Girdaar, 2004

Fukuyama, et. al. 2004

A. MELFO

•

AN OVERCONSTRAINED MODEL

After fine-tune of the SM Higgs mass: 8 parameters left in the heavy Higgs sector

 $m, \alpha, \overline{\alpha}, |\lambda|, |\eta|, \phi = \arg \lambda = -\arg \eta$

$$x = \Re(x) + i\Im(x)$$

n ratio of masses

22

Vevs and masses of all states have form:

$$\sim rac{m}{\lambda} f(x) \ rac{m}{\sqrt{\lambda\eta}} f(x)$$

- variation with parameters quite smooth, with *x* non -trivial



FERMION MASS FITTING

24

• The light Higgs is a combination no longer arbitrary

 $H_{u,d} = r_{u,d}^{10} H_{u,d}^{10} + r_{u,d}^{\overline{126}} H_{u,d}^{\overline{126}} + r_{u,d}^{126} H_{u,d}^{126} + r_{u,d}^{210} H_{u,d}^{210}$

• $r_{u,d}^{\mathbf{I}}$ known functions of the parameters

• Assume type II see-saw

$$m_{\nu} = y_{126} v_{\Delta} \qquad v_{\Delta} = \frac{(\alpha r_u^{10} + \sqrt{6} \eta r_u^{126}) r_u^{210}}{m_{\Delta}}$$

neutrino mass depends on the same parameters

TROUBLE FOR TYPE II SEE-SAW

Some relations among fermion masses depend only on *x*

$$M_u = \frac{N_u}{N_d} \tan \beta \times [M_d + \xi(x)(M_d - M_e)]$$

Define the ratio $R = |1 + 1/\xi(x)|$

then R > 1 from trace identities

25

Write type-II mass as

$$m_{II} = \frac{v^2}{M_x} \times \frac{\sin^2 \beta}{\cos \beta} \times \alpha \sqrt{\frac{|\lambda|}{|\eta|}} \times \frac{M_d - M_e}{v} \times \frac{N_u^2}{N_d} \xi_{II}(x)$$

 \longrightarrow then $\xi_{II}(x)$ must give $10^2 - 10^3$





GENERAL ANALYSIS (TYPE I AND II)

Bertolini, Frigerio, Malinsky, 2005-2006

Aulakh,Garg, Girdaar, 2005-2006

Mohapatra, Goh, Ng, Dutta, Mimura...

• Do the complete fit with all fermion masses and all parameters

Babu, Macesanu Wang, Yang

28

- Parameter space for type I and type II getting smaller
- Include unification constrains, threshold effects
 - even worse

too small neutrino mass: model seems to be ruled out !

WHAT TO DO?

Add more Yukawa couplings: 120

Aulakh, 2005-2006

29

 $D_{120} = (\overline{10}, 1, 1) + (10, 1, 1) + (6, 3, 1) + (6, 1, 3) + (1, 2, 2) + (15, 2, 2)$ (another 10 or 126 cannot help)

- No SM singlets: symmetry breaking is the same
- Antisymmetric: only 3 complex Yukawa couplings more
- Two doublets mix through:

 $c_1 D_{120} H_{10} \Phi_{210} + c_2 D_{120} \Sigma_{126} \Phi_{210} + c_3 D_{120} \overline{\Sigma}_{126} \Phi_{210}$

• More parameters in the superpotential

 $m_D, c_1, c_2, c_3, y_{120}$ 26 + 13 = 39

OR: CHANGE THE HIGGS SECTOR

Alternate model: 54 + 45 instead of 210

- $W = m_H H^2 + m_S S^2 + m_A A^2 + m_\Sigma \Sigma \overline{\Sigma} + \eta A \Sigma \overline{\Sigma} \qquad A:45$ $+ \lambda_H H^2 S + \lambda_S S^3 + \lambda_A A^2 S + \lambda_\Sigma \Sigma^2 S + \lambda_{\overline{\Sigma}} \overline{\Sigma}^2 S \qquad \underline{\Sigma}:126$ H:10
 - 29 real parameters

(compare to 26 in minimal model with 210)

- see-saw of type I and II
- 10 + 126 but...
 - they do not mix light Higgs is only 10

gives wrong fermion masses

30

S:54

	Ŭ
54 + 45 WITH ADDED 120: all doublets mix	
$DAH + DA\Sigma + DA\overline{\Sigma}$	D : 120 H : 10
COMPARE WITH	A:45
MINIMAL MODEL	$\Sigma: 126$
• once you have to enlarge the Yukawa sector, almost same number of parameters: 41 vs. 39	

- smaller representations
- find symmetry breaking and mass spectrum

Ramírez, A.M., in prep.

SYMMETRY BREAKING

$$\langle (1,1,1)_{\mathbf{54}} \rangle \equiv s = \frac{m_A}{\lambda_A} x$$
$$\langle (1,1,15)_{\mathbf{45}} \rangle \equiv a = \frac{m_\Sigma}{\eta_A} \frac{6x-1}{1-2x}$$
$$\langle (1,3,1)_{\mathbf{45}} \rangle \equiv b = \frac{m_\Sigma}{\eta_A} \frac{4x+1}{1-2x}$$
$$\langle (1,3,10)_{\mathbf{126}} \rangle \equiv \langle (1,3,\overline{10})_{\overline{\mathbf{126}}} \rangle \equiv \sigma = \sqrt{\frac{m_\Sigma m_A}{\eta_A^2} \frac{(6x-1)(4x+1)}{1-2x}}$$

32

with

$$\left(\frac{2}{5}\frac{m_{\Sigma}}{m_{A}}\frac{\lambda_{A}}{\eta_{A}}\right)^{2}\frac{x-1}{(1-2x)^{2}}-\frac{\lambda_{S}}{\lambda_{A}}x=\frac{m_{S}}{m_{A}}$$

- one-step breaking at $10^{16}GeV$
- can arrange for a Type II see-saw dominance

EXAMPLE:

ARRANGING A LIGHT TRIPLET

RGE in the MSSM

A. ME

$$\ln\left(\frac{M_{GUT}}{M_W}\right) = \left(\frac{1}{\alpha_j} - \frac{1}{\alpha_i}\right)\frac{2\pi}{b_i - b_j} \qquad i = 1, 2, 3$$

33

Suppose the Δ triplet has a mass $< M_{GUT}$ $v_{\Delta} \propto \frac{1}{m_{\Delta}}$ $m_{\nu} = y_{126} v_{\Delta}$ Type II see-saw

other light fields could cancel its contribution to the running

	$SU(3) \times SU(2) \times U(1)$	δb_1	δb_2	δb_3
	$(1,3;\pm 1)$ Δ	9/5	2	0
	$(6,1;\pm 1/3)$	2/5	0	5/2
	$(1,2;\pm 1/2)$	3/10	1/2	0
	Total	5/2	5/2	5/2
LFO				

ALL THESE FIELDS ARE AVAILABLE $\Delta \qquad \begin{array}{c} (S_{133^{-}}, \underline{\Sigma}_{131}) \\ (S_{133^{+}}, \overline{\Sigma}_{\overline{1}31}) \end{array} \qquad \left(\begin{array}{c} m_{S} + 6\lambda_{S}s & -\sqrt{2}\lambda_{\Sigma}\sigma \\ \sqrt{2}\lambda_{\overline{\Sigma}}\overline{\sigma} & m_{\Sigma} - 3\eta a \end{array}\right)$ $\begin{array}{c} (\overline{\Sigma}_{613^0}, C_{611}) \\ (\Sigma_{\overline{6}13^0}, C_{\overline{6}11}) \end{array} \qquad \left(\begin{array}{c} m_{\Sigma} - \eta \, a & -\sqrt{2} \, \beta \, b \\ \sqrt{2} \, \overline{\beta} \, b & m_C - 12 \lambda_C \, s \end{array} \right)$ $(H_{122^{-}}, \Sigma_{122^{-}}, \overline{\Sigma}_{122^{-}}, C_{122^{-}}, C_{1'22^{-}})$ $(H_{122^+}, \overline{\Sigma}_{122^+}, \Sigma_{122^+}, C_{122^+}, C_{1'22^+})$ $\begin{pmatrix} m_H + 3\lambda_H s & 0 & 0 & \alpha b & -\sqrt{3} \alpha a \\ 0 & m_{\Sigma} - \eta b & -5\lambda_{\Sigma} s & \sqrt{\frac{3}{2}} \overline{\beta} a & -\frac{1}{\sqrt{2}} \overline{\beta} (2a-b) \\ 0 & -5\lambda_{\Sigma} s & m_{\Sigma} + \eta b & -\sqrt{\frac{3}{2}} \beta a & -\frac{1}{\sqrt{2}} \beta (2a+b) \\ \alpha b & \sqrt{\frac{3}{2}} \beta a & -\sqrt{\frac{3}{2}} \overline{\beta} a & m_C + 18\lambda_C s & 0 \\ \sqrt{3} \alpha a & \frac{1}{\sqrt{2}} \beta (2a-b) & \frac{1}{\sqrt{2}} \overline{\beta} (2a+b) & 0 & m_C - 2\lambda_C s \end{pmatrix}$ 34

- enough free parameters to tune their masses at an intermediate scale
- triplet can be as light as necessary for neutrino mass without affecting unification constraints

SUMMARY

35

- SO(10): ideal framework for small neutrino mass
- Models can provide connections between fermion masses and mixings, for example
 - $b \tau$ unification \iff large θ_{atm} (10+126)
 - large neutrino small quark mixings (120+126)
 - ▶ large θ_{atm} → degenerate neutrinos (120+126)
- Non supersymmetric models are alive and well
- Minimal SUSY GUT is in trouble
 * lack of intermediate scales
- Next-to-minimal SUSY GUT may not be predictive ...
 * but work is in progress